



Akbar Shabanloui

EGU 2007 Session GNSS (G6) Wien, Österreich 16 April 2007







LEO Missions (Earth Explorers)





✓ Precise LEO orbits can be used to recover the gravity field of the Earth by the POD method

✓ Analysis of altimetry observations requires precise orbits

✓ Atmosphere sounding requires precise positions of the LEO satellites

✓ GNSS (GPS, GLONASS, GALILEO) methods play an important role in POD in addition to classical methods (e.g. SLR...)







Precise Orbit Determination (POD)









✓ Geometrical orbit determination Only geometrical observations are used, no force models and no constraints; pointwise representation

 ✓ Kinematical orbit determination
 Geometrical, kinematical observations are used, no force models used;
 representation by kinematical functions

✓ Dynamical orbit determination Geometrical, kinematical & dynamical observations are used, force models necessary; pointwise representation or representation by functions







Geometrical precise orbit determination



Observations:

Only geometrical SST observations between GPS satellites and LEQ satellite in MULL NAMA case of one LEO satellite,

Gauss-Markov model

Unknowns

Model:

LEO positions at the observed epochs (point wise).







Kinematical precise orbit determination

7

Observations:

Geometrical SST observations between GPS satellites and LEO satellite in case of one LEO satellite, LEO representation model with respect to physical background

Model:

Gauss-Markov model,

Unknowns:

LEO short arc representation parameters (continuous)







Dynamical precise orbit determination

Observations:

Geometrical SST observations between GPS satellites and LEO satellite in case of one LEO satellite, LEO representation model with respect to physical background, Earth gravity field, dynamical observations

Model:

Gauss-Markov model

nwn

LEO boundary positions, (continuous)







Processing concepts

















Step by step presentation of short arcs of LEO





selected a-priori

$$\mathbf{r}(t) = \mathbf{r}_a \frac{\sin((1-\tau)N)}{\sin(N)} + \mathbf{r}_e \frac{\sin(\tau N)}{\sin(N)} + \mathbf{C}^T \mathbf{P}(\tau) + \sum_{\nu=1}^n \overline{\mathbf{d}}_{\nu} \sin(\nu \pi \tau)$$

all coefficients can be estimated by a Gauss-Markov model

Advantage:

the kinematical orbits and another kinematical parameters can be derived directly from estimated LEO short arc parameters.















cut-off angle 15°

simple data processing & S/N filtering

elevation weighting





Sequential time differenced carrier phase observations can be written as:

$$\Delta \Phi_{r}^{s}(t_{1}, t_{2}) = \left| \mathbf{r}^{s}(t_{2}) - \mathbf{r}_{r}(t_{2}) \right| - \left| \mathbf{r}^{s}(t_{1}) - \mathbf{r}_{r}(t_{1}) \right| + e_{\Delta \Phi_{3}}$$

$$\Delta \Phi_{r,3}^{s}(t_{1}, t_{2}) = \left| \mathbf{r}^{s}(t_{2}) - (\mathbf{r}_{a} \cdot \frac{\sin((1 - \tau_{2})N)}{\sin(N)} + \mathbf{r}_{e} \cdot \frac{\sin(\tau_{2}N)}{\sin(N)} + \mathbf{C}_{3\times 4} \mathbf{P}(\tau_{2}) + \sum_{f=1}^{n} \mathbf{d}_{f} \cdot \sin(\pi f \tau_{2}) \right| - \left| \mathbf{r}^{s}(t_{1}) - (\mathbf{r}_{a} \cdot \frac{\sin((1 - \tau_{1})N)}{\sin(N)} + \mathbf{r}_{e} \cdot \frac{\sin(\tau_{1}N)}{\sin(N)} + \mathbf{C}_{3\times 4} \mathbf{P}(\tau_{1}) + \sum_{f=1}^{n} \mathbf{d}_{f} \cdot \sin(\pi f \tau_{1}) \right| + e_{\Delta \Phi_{3}}$$







The Gauss-Markov model for one epoch,

$$\begin{pmatrix} \Delta \Delta \Phi_{r}^{s_{1}}(t_{1},t_{2}) \\ \vdots \\ \Delta \Delta \Phi_{r}^{s_{j}}(t_{1},t_{2}) \\ \vdots \\ \Delta \Delta \Phi_{r}^{s_{j}}(t_{1},t_{2}) \\ \vdots \\ \Delta \Delta \Phi_{r}^{s_{m}}(t_{1},t_{2}) \end{pmatrix} = \begin{pmatrix} \mathbf{A}_{\mathbf{r}_{r}}^{s_{1}}(t_{2})\mathbf{A}_{\mathbf{X}_{r}}^{\mathbf{r}_{r}}(t_{2}) - \mathbf{A}_{\mathbf{r}_{r}}^{s_{j}}(t_{1})\mathbf{A}_{\mathbf{X}_{r}}^{\mathbf{r}_{r}}(t_{1}) \\ \vdots \\ \mathbf{A}_{\mathbf{T}_{r}}^{s_{j}}(t_{2})\mathbf{A}_{\mathbf{X}_{r}}^{\mathbf{r}_{r}}(t_{2}) - \mathbf{A}_{\mathbf{r}_{r}}^{s_{j}}(t_{1})\mathbf{A}_{\mathbf{X}_{r}}^{\mathbf{r}_{r}}(t_{1}) \\ \vdots \\ \mathbf{A}_{\mathbf{T}_{r}}^{s_{m}}(t_{2})\mathbf{A}_{\mathbf{X}_{r}}^{\mathbf{r}_{r}}(t_{2}) - \mathbf{A}_{\mathbf{r}_{r}}^{s_{m}}(t_{1})\mathbf{A}_{\mathbf{X}_{r}}^{\mathbf{r}_{r}}(t_{1}) \end{pmatrix} [\mathbf{X}_{r} - \mathbf{X}_{r}^{0}] \\ \mathbf{X}_{r} = \begin{pmatrix} x_{a} \quad y_{a} \quad z_{a} \quad \cdots \quad z_{e} \quad c_{11} \quad c_{12} \quad c_{13} \quad c_{14} \quad \cdots \quad c_{34} \quad \overline{d}_{1,x} \quad \overline{d}_{1,y} \quad \overline{d}_{1,z} \quad \cdots \quad \overline{d}_{n,z} \end{pmatrix}_{\mathbf{6}+\mathbf{12}+\mathbf{3}n}^{\mathbf{T}}$$









- NA BEEL
- Unknows: boundary values, polynomial coefficients, amplitudes of Fourier series.
- Solutions: Gauss-Markov model
- Convergence & accuracy: after a few iterations, ~ cm.

 LEO short arc orbit is continuous and velocities and another kinematical parameters can be derived from the estimated parameters.







Kinematical short arc POD-simulated



30 minutes of CHAMP satellite [2000 07 15 15h 10m - 15h 40m]







Kinematical short arc POD sequential time differenced carrier phase

Difference plot between estimated short arc absolute positions with observation precision=0.01 m & given positions













$$\Phi_{r,i}^{s}(t) = \left| \mathbf{R}_{z}(\boldsymbol{\omega}_{e}.\boldsymbol{\varepsilon}_{r}^{s})\mathbf{r}^{s}(t-\boldsymbol{\varepsilon}_{r}^{s}) - \mathbf{r}_{r}(t) \right| + c \left[dt^{s}(t-\boldsymbol{\varepsilon}_{r}^{s}) - dt_{r}(t) \right] + \lambda N_{r}^{s} + I_{i}^{r}(t) + d_{O}^{s}(t) + d_{R}^{r}(t) - d_{R}^{s}(t) + d_{C,i}^{r}(t) + d_{V,i}^{r}(t) + d_{M,P_{i}}(t) + e_{P_{i}}$$

GPS, LEO indices,

Travelling time between GPS & LEO,

 $\mathbf{r}^{s}(t-\boldsymbol{\mathcal{E}}_{r}^{s}), dt^{s}(t-\boldsymbol{\mathcal{E}}_{r}^{s})$

GPS position, clock offset at sending time,

 $\mathbf{r}_{r}(t), dt_{r}(t)$

LEO position, clock offset at receiving time



S, r

 \mathcal{E}_r^s





universitätbo

Carrier phase GPS SST observation...



Carrier phase ionosphere-free observation at epochs (1,2) $\Phi_{r,3}^{s}(t_{1}) = |\mathbf{R}_{z}(\omega_{e}\varepsilon_{1})\mathbf{r}^{s}(t_{1}-\varepsilon_{1})-\mathbf{r}_{r}(t_{1})| + \lambda_{3}N_{r,3}^{s} + c[dt^{s}(t_{1}-\varepsilon_{1})-dt_{r}(t_{1})] + d_{O}^{s}(t_{1}) + d_{R}^{r}(t_{1}) - d_{R}^{s}(t_{1}) + d_{C,3}^{r}(t_{1}) + d_{V,3}^{r}(t_{1}) + d_{M,\Phi_{3}}(t_{1}) + e_{\Phi_{3}}$ $\Phi_{r,3}^{s}(t_{2}) = |\mathbf{R}_{z}(\omega_{e}\varepsilon_{2})\mathbf{r}^{s}(t_{2}-\varepsilon_{2})-\mathbf{r}_{r}(t_{2})| + \lambda_{3}N_{r,3}^{s} + c[dt^{s}(t_{2}-\varepsilon_{2})-dt_{r}(t_{2})] + d_{O}^{s}(t_{2}) + d_{R}^{r}(t_{2}) - d_{R}^{s}(t_{2}) + d_{C,3}^{r}(t_{2}) + d_{V,3}^{r}(t_{2}) + d_{M,\Phi_{3}}(t_{2}) + e_{\Phi_{3}}$

sequential time difference ionosphere-free carrier phase observation between epochs (1,2)

 $\Delta \Phi_{r,3}^{s}(t_1,t_2) = \left| \mathbf{R}_{z}(\boldsymbol{\omega}_{e}\boldsymbol{\varepsilon}_{2}) \cdot \mathbf{r}^{s}(t_2 - \boldsymbol{\varepsilon}_{2}) - \mathbf{r}_{r}(t_2) \right| - \left| \mathbf{R}_{z}(\boldsymbol{\omega}_{e}\boldsymbol{\varepsilon}_{1}) \cdot \mathbf{r}^{s}(t_1 - \boldsymbol{\varepsilon}_{1}) - \mathbf{r}_{r}(t_1) \right| - \mathbf{v}_{z}(t_1 - \boldsymbol{\varepsilon}_{1}) \cdot \mathbf{r}^{s}(t_1 - \boldsymbol{\varepsilon}_{1}) - \mathbf{v}_{z}(t_1) \right|$







Ionosphere free sequential time differenced carrier phase observations can be written as:

$$\Delta \Phi_{r,3}^{s}(t_{1},t_{2}) = \left| \mathbf{R}_{z}(\boldsymbol{\omega}_{e}\boldsymbol{\varepsilon}_{2}) \mathbf{r}^{s}(t_{2}-\boldsymbol{\varepsilon}_{2}) - \mathbf{r}_{r}(t_{2}) \right| - \left| \mathbf{R}_{z}(\boldsymbol{\omega}_{e}\boldsymbol{\varepsilon}_{1}) \mathbf{r}^{s}(t_{1}-\boldsymbol{\varepsilon}_{1}) - \mathbf{r}_{r}(t_{1}) \right|$$

$$c\Delta dt_{r}(t_{1},t_{2}) + e_{\Delta \Phi_{3}}$$

$$\Delta \Phi_{r,3}^{s}(t_{1},t_{2}) = \left| \mathbf{R}_{z}(\boldsymbol{\omega}_{e}\boldsymbol{\varepsilon}_{2}) \mathbf{r}^{s}(t_{2}-\boldsymbol{\varepsilon}_{2}) - (\mathbf{r}_{a}.\frac{\sin((1-\tau_{2})N)}{\sin(N)} + \mathbf{r}_{e}.\frac{\sin(\tau_{2}N)}{\sin(N)} + \mathbf{C}_{3\times 4}\mathbf{P}(\tau_{2}) + \sum_{f=1}^{n} \overline{\mathbf{d}}_{f}\sin(\pi f \tau_{2})) \right| - \left| \mathbf{R}_{z}(\boldsymbol{\omega}_{e}\boldsymbol{\varepsilon}_{1}) \mathbf{r}^{s}(t_{1}-\boldsymbol{\varepsilon}_{1}) - (\mathbf{r}_{a}.\frac{\sin((1-\tau_{1})N)}{\sin(N)} + \mathbf{r}_{e}.\frac{\sin(\tau_{1}N)}{\sin(N)} + \mathbf{C}_{3\times 4}\mathbf{P}(\tau_{1}) + \sum_{f=1}^{n} \overline{\mathbf{d}}_{f}\sin(\pi f \tau_{1})) \right| - c\Delta dt_{r}(t_{1},t_{2}) + e_{\Delta \Phi_{3}}$$
universitätioon





The Gauss-Markov model for one epoch,

$$\begin{pmatrix} \Delta \Delta \Phi_{r}^{s_{1}}(t_{1},t_{2}) \\ \vdots \\ \Delta \Delta \Phi_{r}^{s_{j}}(t_{1},t_{2}) \\ \vdots \\ \Delta \Delta \Phi_{r}^{s_{j}}(t_{1},t_{2}) \end{pmatrix} = \begin{pmatrix} \mathbf{A}_{r_{r}}^{s_{i}}(t_{2})\mathbf{A}_{\mathbf{X}_{r}}^{\mathbf{r}}(t_{2}) - \mathbf{A}_{r_{r}}^{s_{j}}(t_{1})\mathbf{A}_{\mathbf{X}_{r}}^{\mathbf{r}}(t_{1}) \\ \vdots \\ \mathbf{A} \Phi_{r}^{s_{j}}(t_{1},t_{2}) \end{pmatrix} = \begin{pmatrix} \mathbf{A}_{r_{r}}^{s_{j}}(t_{2})\mathbf{A}_{\mathbf{X}_{r}}^{\mathbf{r}}(t_{2}) - \mathbf{A}_{r_{r}}^{s_{j}}(t_{1})\mathbf{A}_{\mathbf{X}_{r}}^{\mathbf{r}}(t_{1}) \\ \vdots \\ \mathbf{A} \Delta \Phi_{r}^{s_{m}}(t_{1},t_{2}) \end{pmatrix} = \begin{pmatrix} \mathbf{A}_{r_{r}}^{s_{m}}(t_{2})\mathbf{A}_{\mathbf{X}_{r}}^{\mathbf{r}}(t_{2}) - \mathbf{A}_{r_{r}}^{s_{j}}(t_{1})\mathbf{A}_{\mathbf{X}_{r}}^{\mathbf{r}}(t_{1}) \\ \mathbf{A}_{\mathbf{X}_{r}}^{\mathbf{r}}(t_{2})\mathbf{A}_{\mathbf{X}_{r}}^{\mathbf{r}}(t_{2}) - \mathbf{A}_{r_{r}}^{s_{m}}(t_{1})\mathbf{A}_{\mathbf{X}_{r}}^{\mathbf{r}}(t_{1}) \end{pmatrix} = \begin{bmatrix} \mathbf{A}_{r}^{s_{m}}(t_{1})\mathbf{A}_{\mathbf{X}_{r}}^{\mathbf{r}}(t_{1}) \\ \mathbf{A}_{r_{r}}^{s_{m}}(t_{2})\mathbf{A}_{\mathbf{X}_{r}}^{\mathbf{r}}(t_{2}) - \mathbf{A}_{r_{r}}^{s_{m}}(t_{1})\mathbf{A}_{\mathbf{X}_{r}}^{\mathbf{r}}(t_{1}) \end{bmatrix} = \begin{bmatrix} \mathbf{A}_{r}^{s_{m}}(t_{1})\mathbf{A}_{\mathbf{X}_{r}}^{\mathbf{r}}(t_{1}) \\ \mathbf{A}_{r_{r}}^{s_{m}}(t_{1})\mathbf{A}_{\mathbf{X}_{r}}^{\mathbf{r}}(t_{2}) - \mathbf{A}_{r_{r}}^{s_{m}}(t_{1})\mathbf{A}_{\mathbf{X}_{r}}^{\mathbf{r}}(t_{1}) \end{bmatrix} = \begin{bmatrix} \mathbf{A}_{r}^{s_{m}}(t_{1})\mathbf{A}_{\mathbf{X}_{r}}^{\mathbf{r}}(t_{1}) \\ \mathbf{A}_{r_{r}}^{s_{m}}(t_{2})\mathbf{A}_{\mathbf{X}_{r}}^{\mathbf{r}}(t_{2}) - \mathbf{A}_{r_{r}}^{s_{m}}(t_{1})\mathbf{A}_{\mathbf{X}_{r}}^{\mathbf{r}}(t_{1}) \end{bmatrix} = \begin{bmatrix} \mathbf{A}_{r}^{s_{m}}(t_{1})\mathbf{A}_{r_{r}}^{\mathbf{r}}(t_{1}) \\ \mathbf{A}_{r_{r}}^{s_{m}}(t_{2})\mathbf{A}_{\mathbf{X}_{r}}^{\mathbf{r}}(t_{2}) - \mathbf{A}_{r_{r}}^{s_{m}}(t_{1})\mathbf{A}_{\mathbf{X}_{r}}^{\mathbf{r}}(t_{1}) \end{bmatrix} = \begin{bmatrix} \mathbf{A}_{r}^{s_{m}}(t_{1})\mathbf{A}_{r_{r}}^{\mathbf{r}}(t_{1}) \\ \mathbf{A}_{r}^{s_{m}}(t_{1})\mathbf{A}_{r_{r}}^{\mathbf{r}}(t_{1})\mathbf{A}_{r_{r}}^{\mathbf{r}}(t_{1}) \\ \mathbf{A}_{r_{r}}^{s_{m}}(t_{1})\mathbf{A}_{r_{r}}^{s_{m}}(t_{1})\mathbf{A}_{r_{r}}^{s_{m}}(t_{1})\mathbf{A}_{r_{r}}^{s_{m}}(t_{1}) \end{bmatrix} = \begin{bmatrix} \mathbf{A}_{r}^{s_{m}}(t_{1})\mathbf{A}_{r_{r}}^{s_{m}}(t_{1})\mathbf{A}_{r_{r}}^{s_{m}}(t_{1})\mathbf{A}_{r_{r}}^{s_{m}}(t_{1}) \\ \mathbf{A}_{r_{r}}^{s_{m}}(t_{1})\mathbf{A}_{r_{r}}^{s_{m}}(t_{1})\mathbf{A}_{r_{r}}^{s_{m}}(t_{1})\mathbf{A}_{r_{r}}^{s_{m}}(t_{1})\mathbf{A}_{r_{r}}^{s_{m}}(t_{1}) \end{bmatrix} = \begin{bmatrix} \mathbf{A}_{r}^{s_{m}}(t_{1})\mathbf{A}_{r_{r}}^{s_{m}}(t_{1})\mathbf{A}_{r_{r}}^{s_{m}}(t_{1})\mathbf{A}_{r_{r}}^{s_{m}}(t_{1})\mathbf{A}_{r_{r}}^{s_{m}}(t_{1})\mathbf{A}_{r_{r}}^{s_{m}}(t_{1})\mathbf{A}_{r_{$$





- Initial values: unknowns initial values can be derived from code estimated positions at first step,
- Unknows: boundary values, polynomial coefficients, amplitudes of Fourier series,
- Solutions: Gauss-Markov model,
- **Convergence & accuracy:** after ~a few iterations, ~ cm.

 LEO short arc orbit is continuous and velocities and another kinematical parameters can be derived from the estimated LEO short arc parameters.



Kinematical short arc POD with sequential time

Difference plot between absolute positions with carrier phase observation precision =0.01 m & given GFZ CHAMP PSO orbit





- From sequential differenced carrier phase SST observations, LEO absolute positions and clock offsets can be estimated at every epoch with (or even without) enough number of GPS satellites (4?) and good satellite geometry
- An accuracy of cm can be expected for the sequnetial time diffenced carrier phase SST data processing, but DOP! isn't crucial
- The resulting LEO orbit is given continuous (without gaps)
 & smoother than geometrical POD
- Kinematical LEO orbit can be used to recover the Earth's gravity field with the POD recovery concept. (kinematical parameters can be derived analytically)









Thank you for your

attentions





