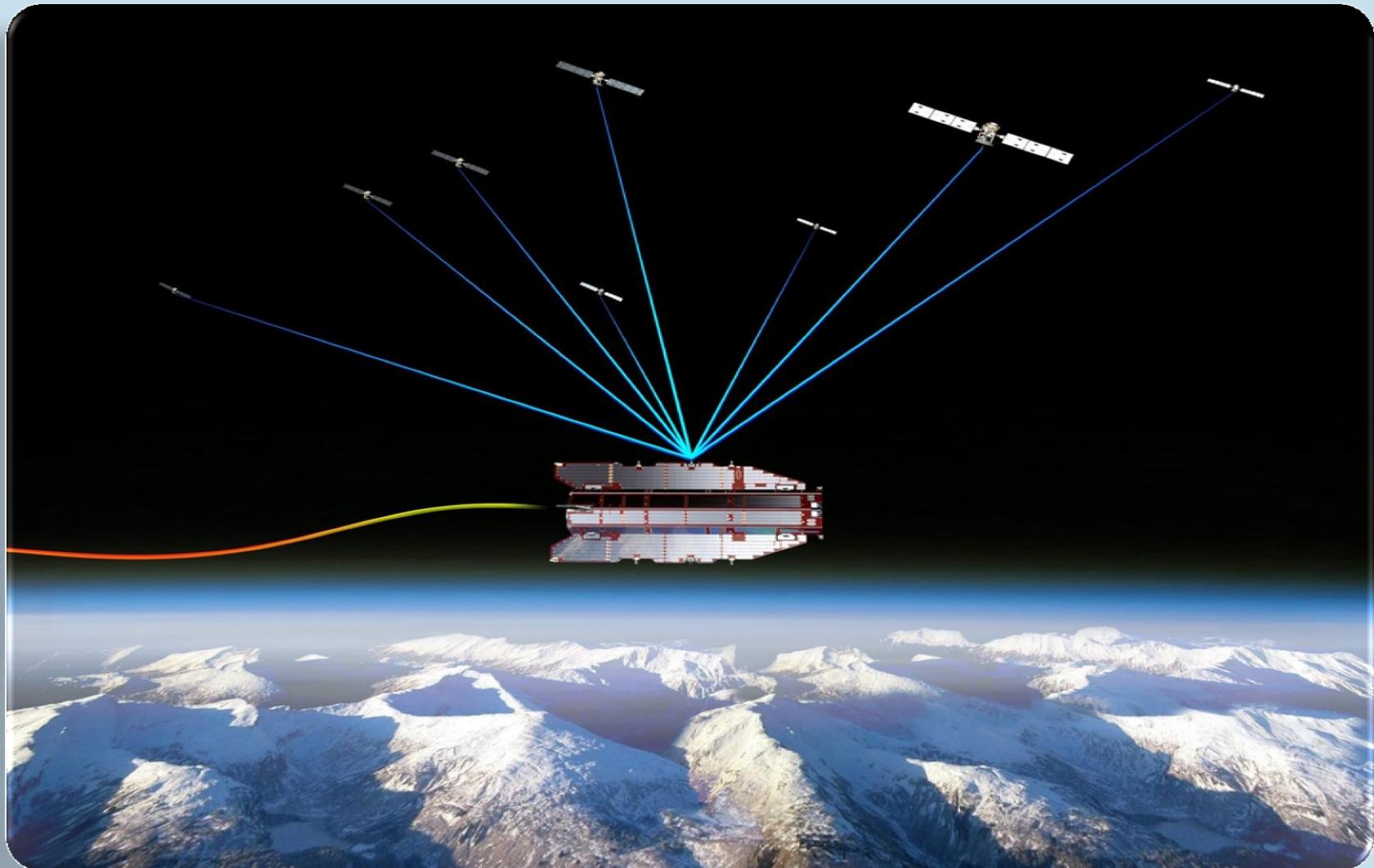


# From pure kinematical to reduced kinematical LEO orbit determination

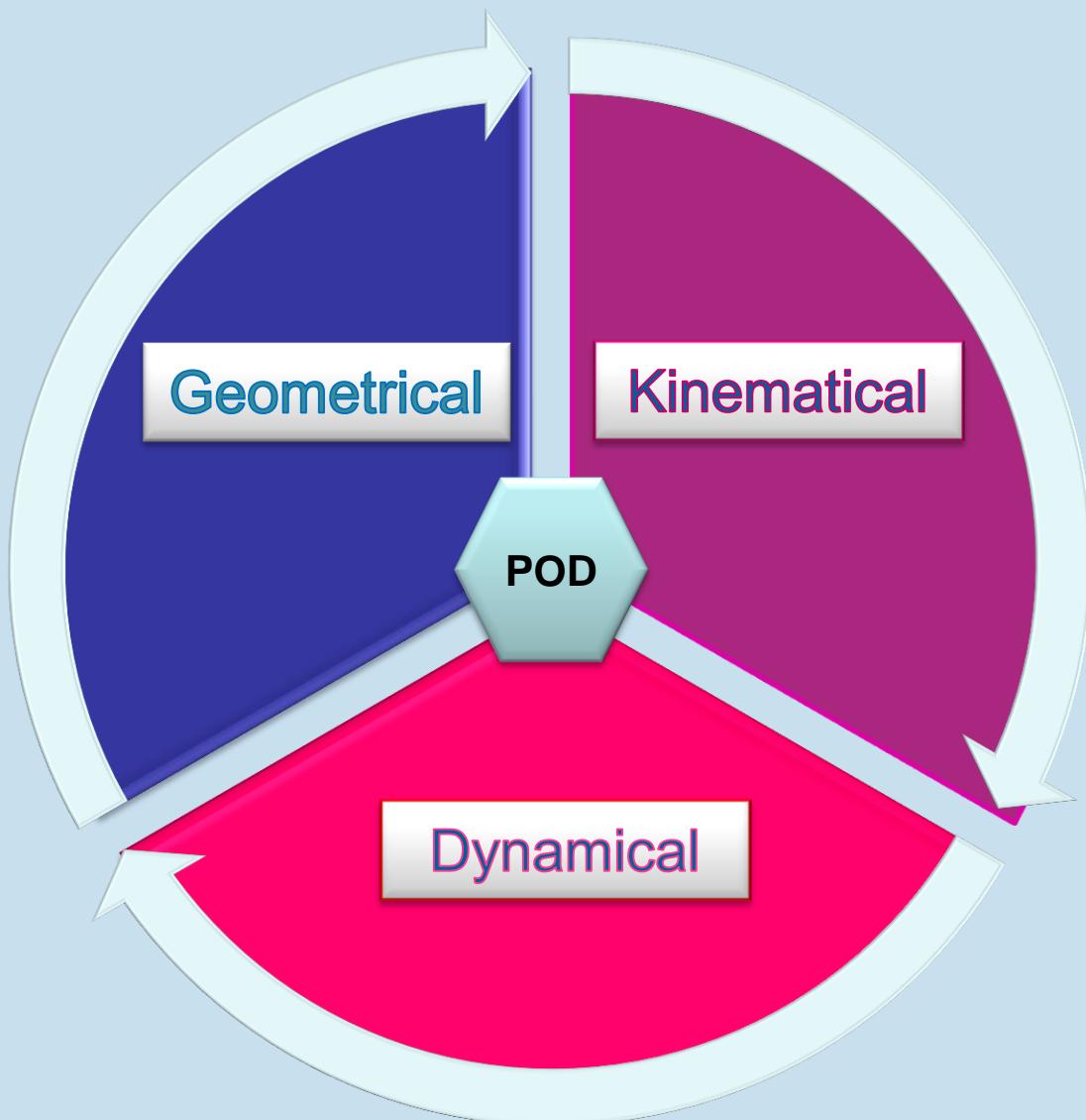
Akbar Shabanloui

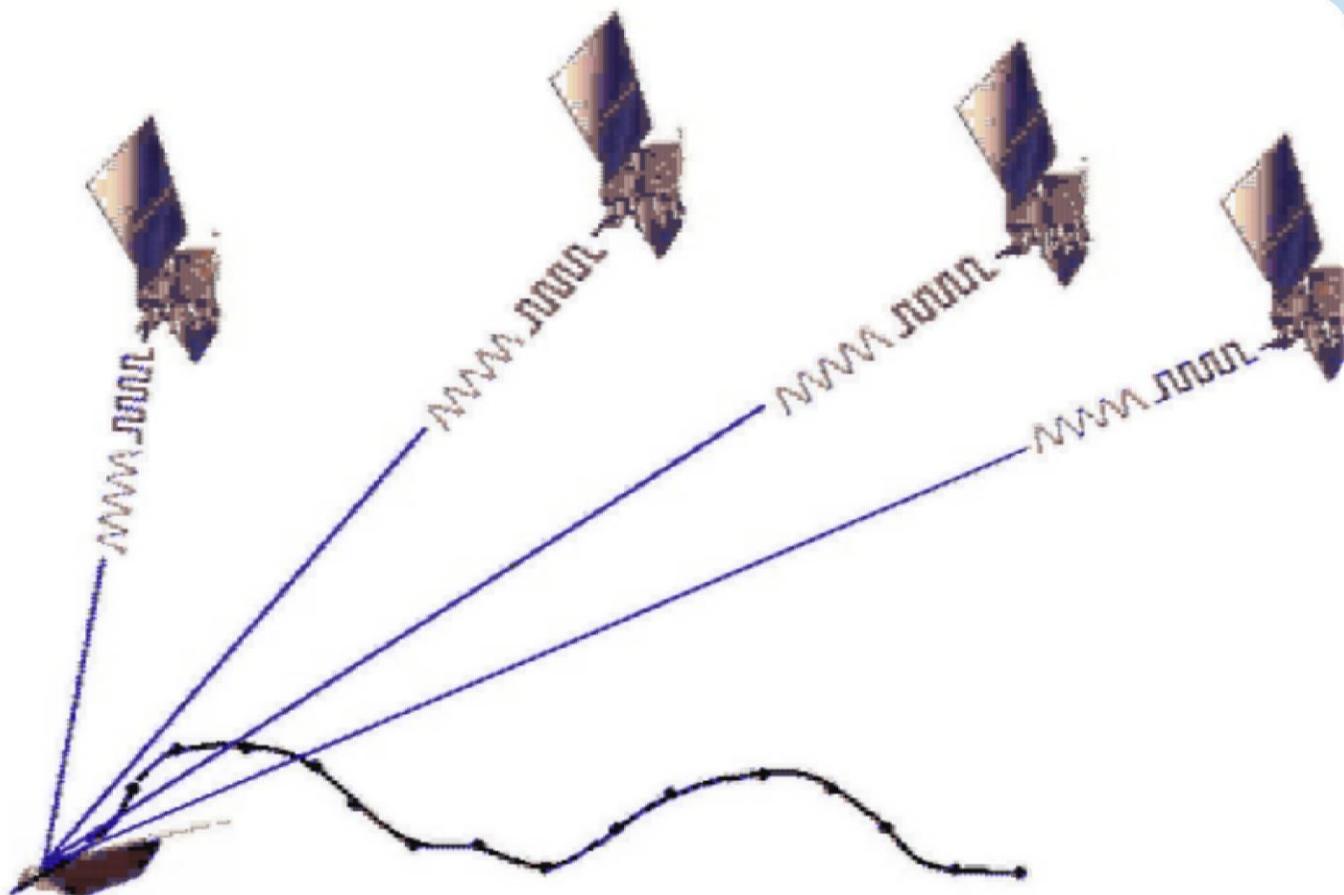
Session 4  
Geodätische Woche 2009  
23 Sep. 2009, Karlsruhe, Deutschland

# Precise Orbit Determination (POD)



Credit: European Space Agency





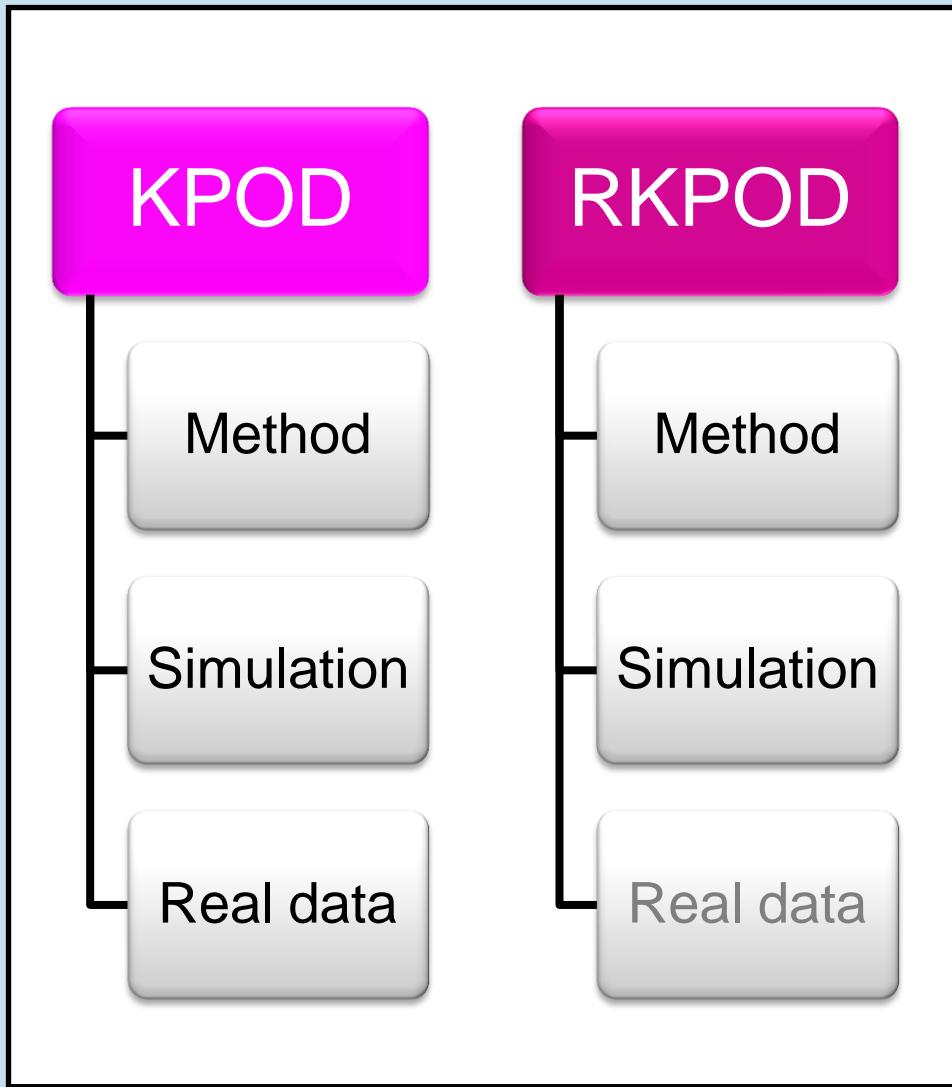
GPOD

KPOD

RKPOD

DPOD

5



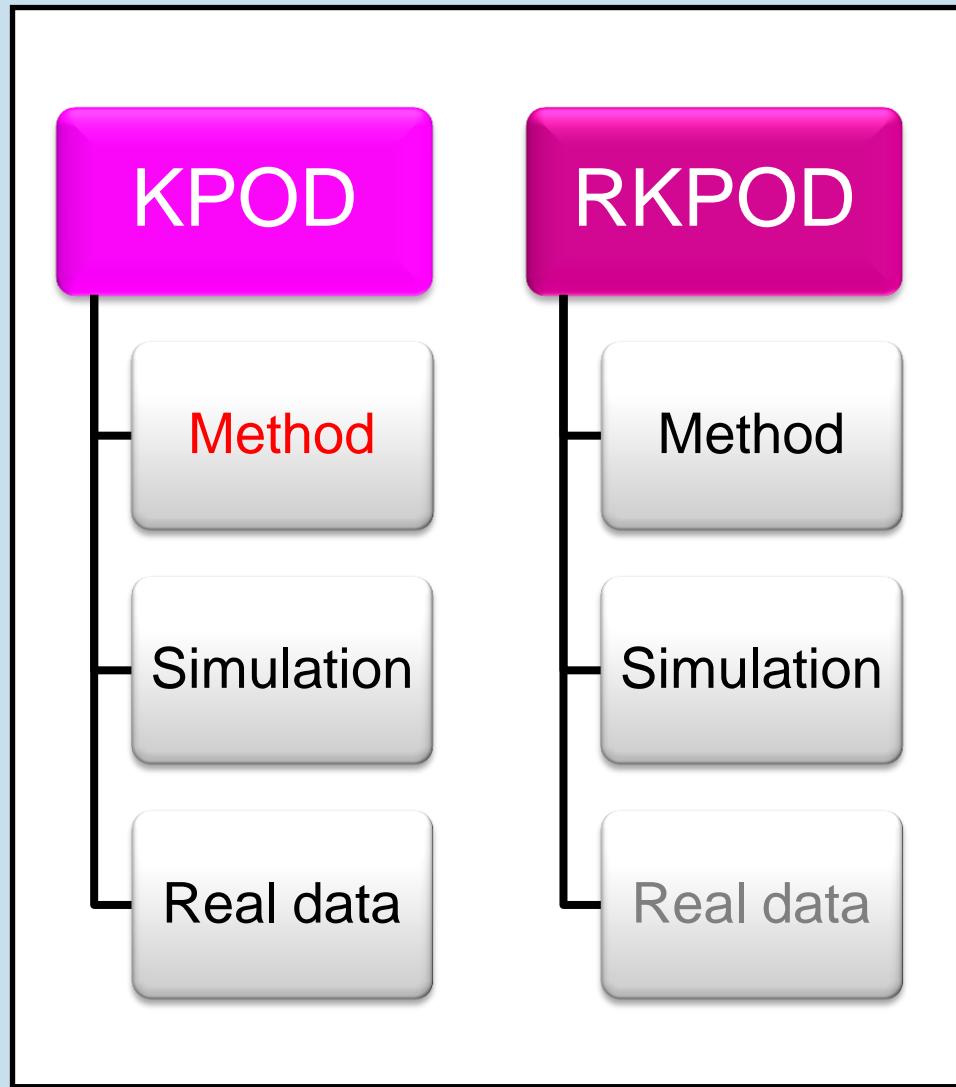
GPOD

KPOD

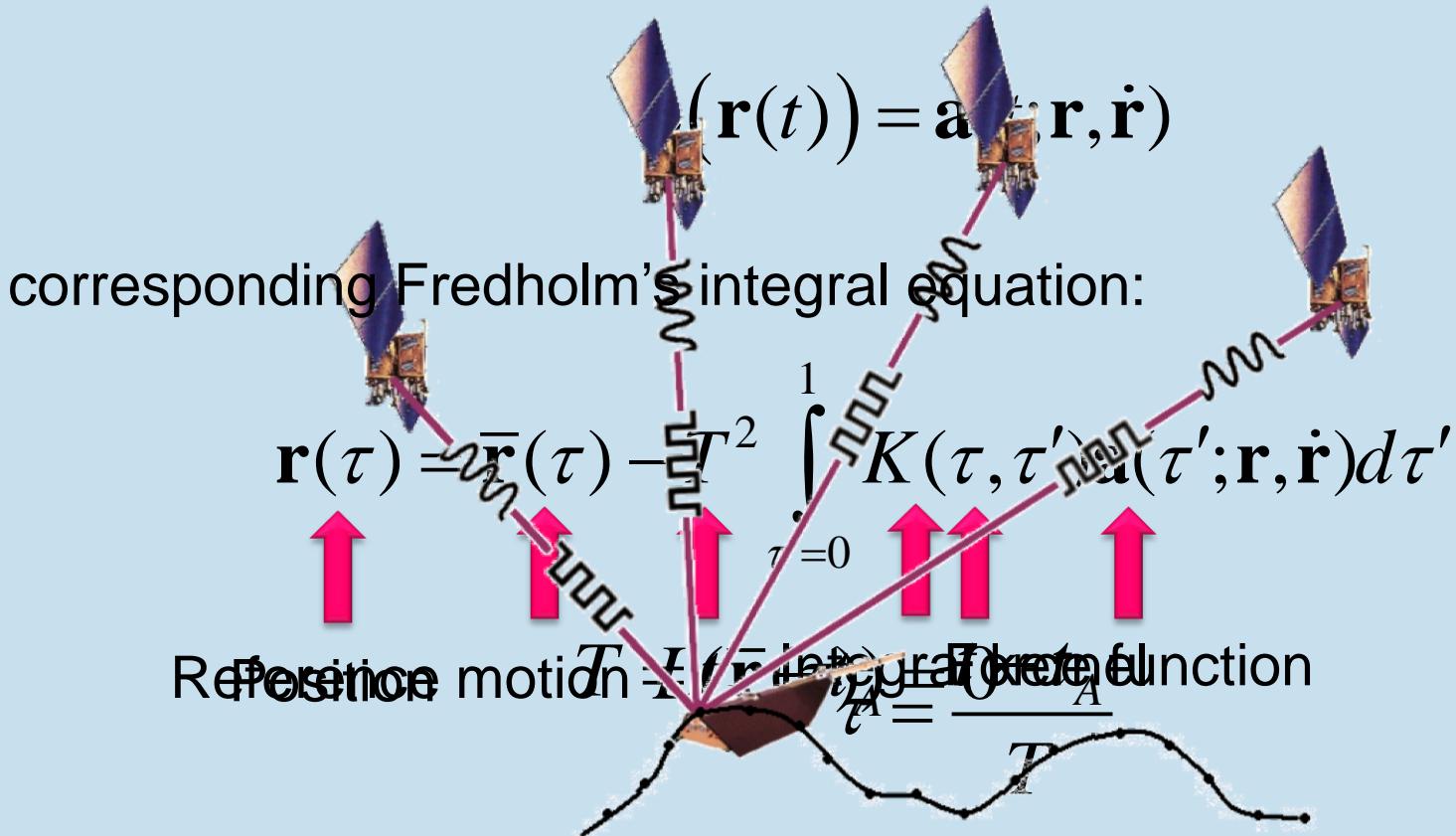
RKPOD

DPOD

6



Equation of motion w.r.t self-adjoint differential operator:



A satellite short arc:

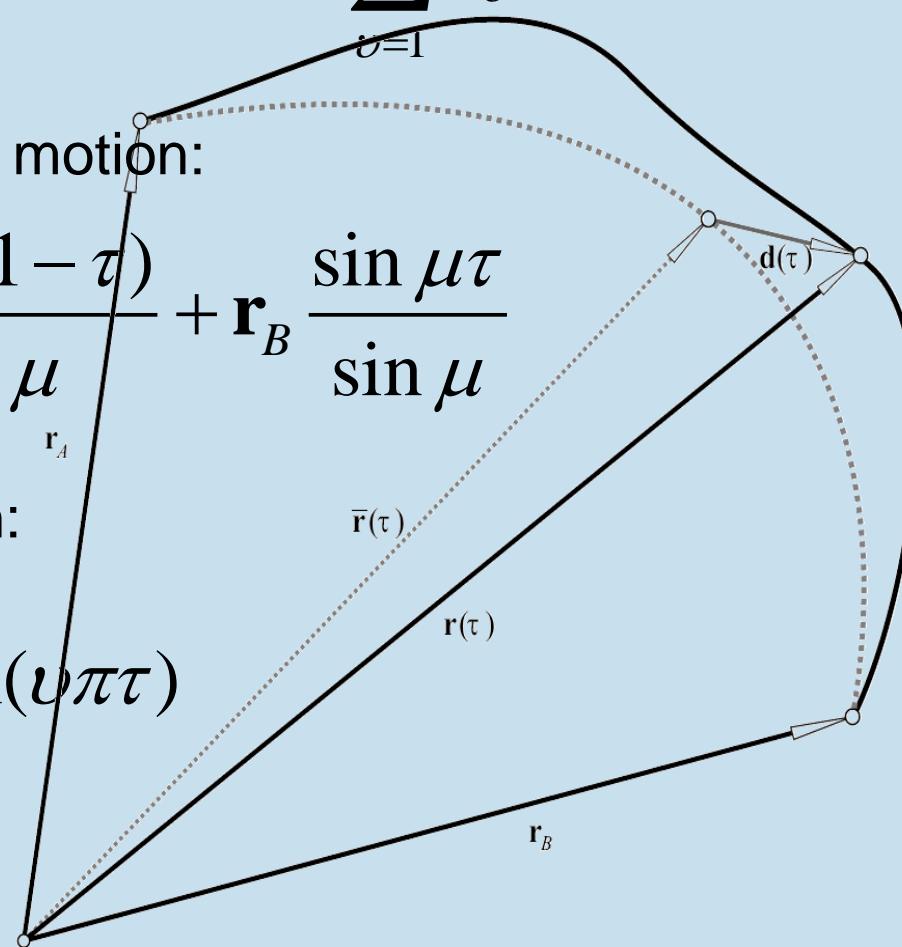
$$\mathbf{r}(\tau) = \bar{\mathbf{r}}(\tau) + \mathbf{d}(\tau) = \bar{\mathbf{r}}(\tau) + \sum_{v=1}^{\infty} \mathbf{d}_v \sin(v\pi\tau)$$

elliptical reference motion:

$$\bar{\mathbf{r}}(\tau) = \mathbf{r}_A \frac{\sin \mu(1-\tau)}{\sin \mu} + \mathbf{r}_B \frac{\sin \mu\tau}{\sin \mu}$$

difference function:

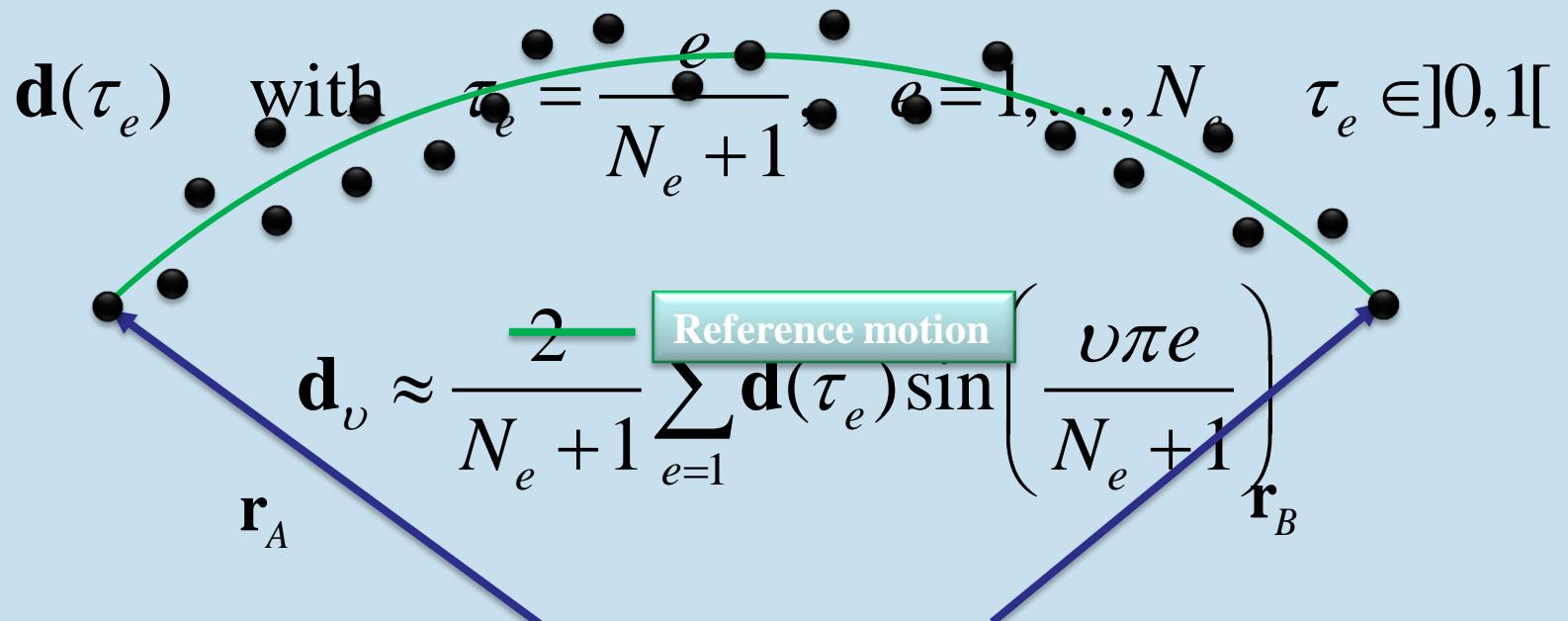
$$\mathbf{d}(\tau) = \sum_{v=1}^{\infty} \mathbf{d}_v \sin(v\pi\tau)$$

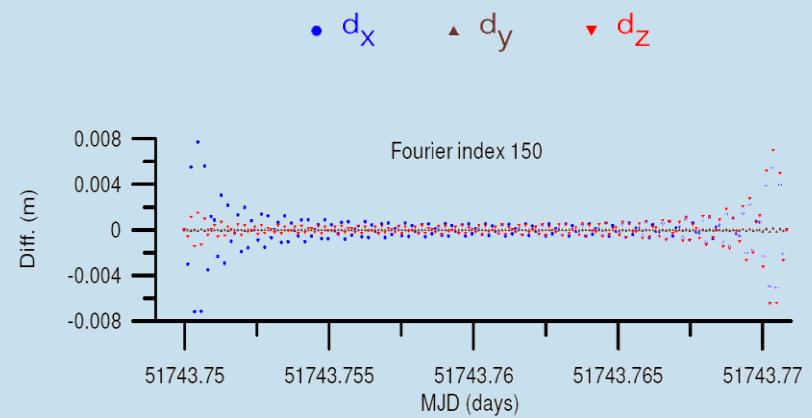
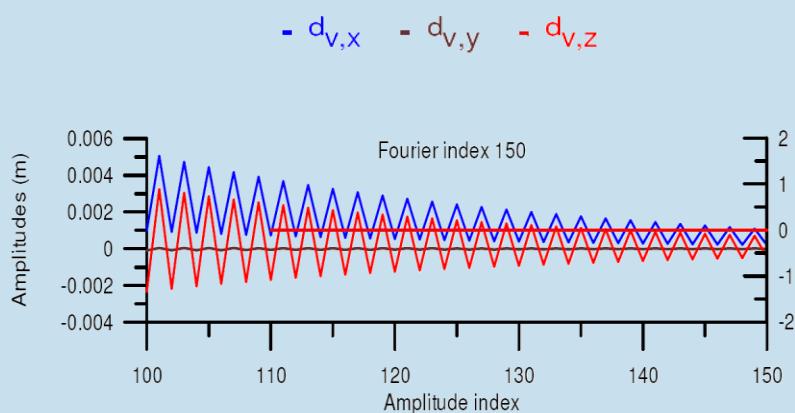


A satellite short arc can be represented:

$$\mathbf{r}(\tau) \approx \mathbf{r}_A + \dot{\mathbf{r}}(\tau) (\tau - \tau_A) + \ddot{\mathbf{r}}(\tau) \frac{(\tau - \tau_A)^2}{2!} + \ddots + \sum_{v=1}^{\infty} \sum_{\nu=1}^{\infty} \mathbf{d}_v \sin(v\pi\nu(\tau - \tau_A))$$

$$\mathbf{d}_v = 2 \int_{\tau'=0}^1 \mathbf{d}(\tau) \sin(v\pi\tau') d\tau'$$





Amplitudes

Remainders

A satellite short arc can be represented:

$$\mathbf{r}(\tau) = \bar{\mathbf{r}}(\tau) + \sum_{v=1}^{\infty} \mathbf{d}_v \sin(v\pi\tau)$$

with Fourier amplitudes:

$$\mathbf{d}_v = 2 \int_{\tau'=0}^1 \mathbf{d}(\tau) \sin(v\pi\tau') d\tau'$$

Fourier series amplitudes:

$$\begin{aligned} \mathbf{d}_v &= \sum_{j=1}^J \frac{2(-1)^{j+1}}{(\nu\pi)^{2j+1}} [(-1)^v \mathbf{d}^{[2j]}(1) - \mathbf{d}^{[2j]}(0)] + \\ &\quad + \beta \frac{2}{(\nu\pi)^{2J+1}} \int_{\tau'=0}^1 \mathbf{d}^{[2J+2]}(\tau') \sin(\nu\pi\tau') d\tau' \end{aligned}$$

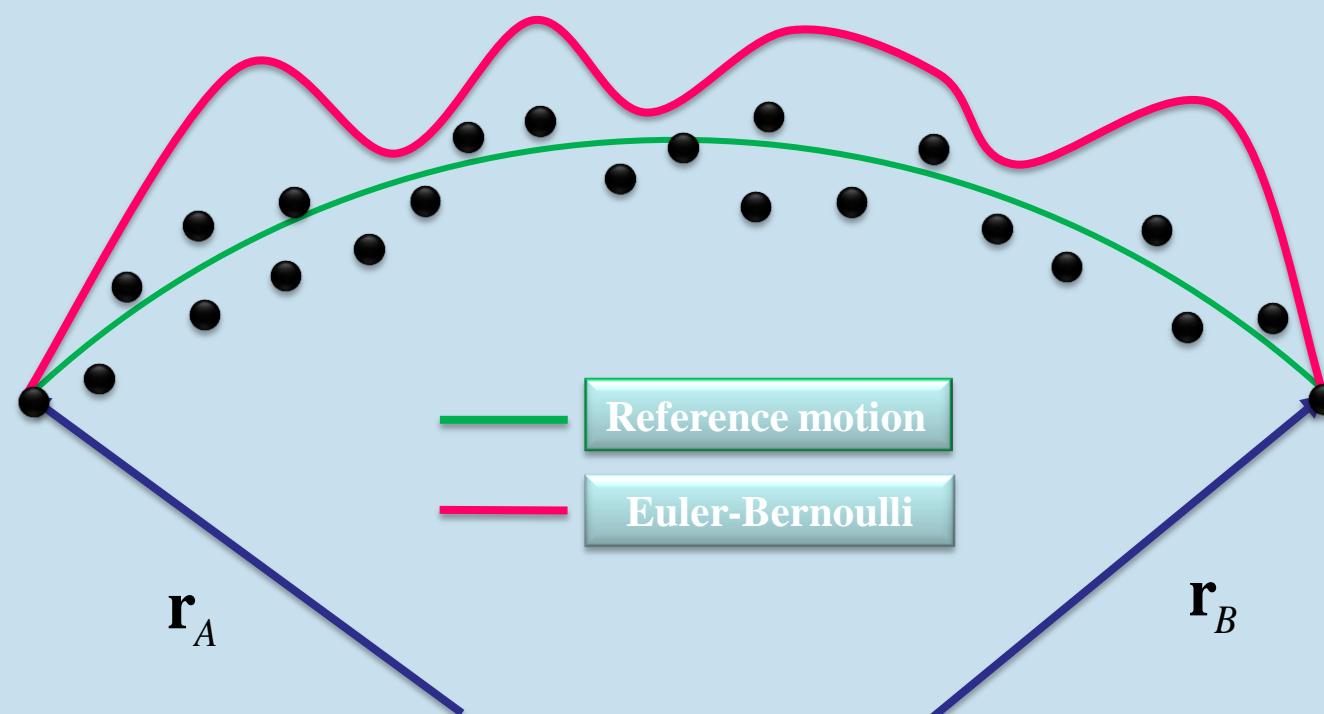
$$\begin{aligned}\mathbf{d}_F^\infty &= \mathbf{d}(\tau) = \sum_{v=1}^{\infty} \mathbf{d}_v \sin(v\pi\tau) = \\ &= \sum_{j=1}^{\infty} \mathbf{e}_{2j} E_{2j}(\tau) + \sum_{j=1}^{\infty} \mathbf{b}_{2j+1} B_{2j+1}(\tau) = \mathbf{d}_P^\infty\end{aligned}$$

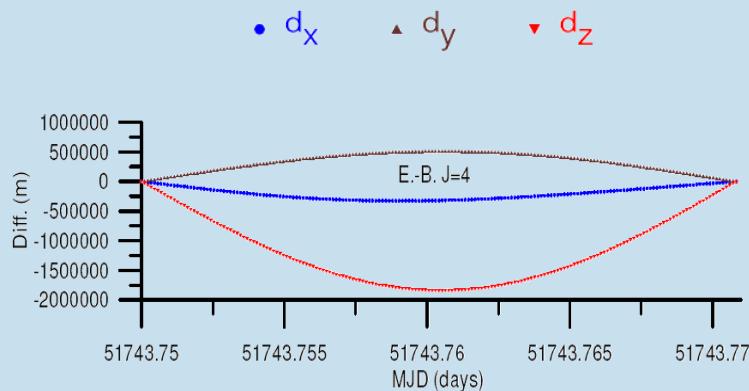
A satellite short arc can be represented:

$$\sum_{v=1}^{\infty} \mathbf{d}_v \sin(v\pi\tau) = \sum_{j=1}^{\infty} \mathbf{e}_{2j} E_{2j}(\tau) + \sum_{j=1}^{\infty} \mathbf{b}_{2j+1} B_{2j+1}(\tau)$$

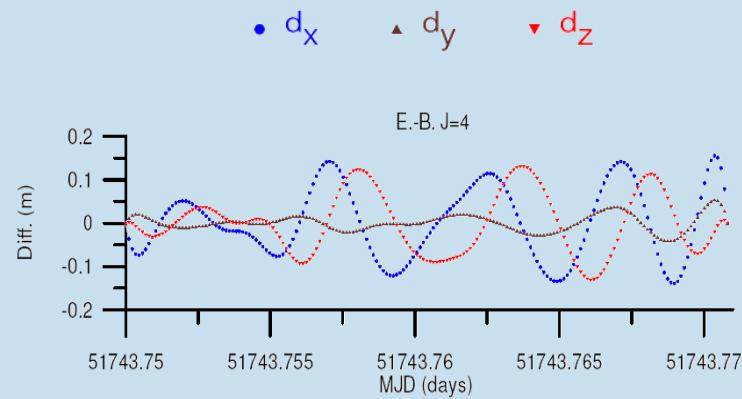
A satellite short arc can be represented with the Euler-Bernoulli term up to degree  $J$  as:

$$\mathbf{r}(\tau) - \bar{\mathbf{r}}(\tau) = \mathbf{d}(\tau) \approx \sum_{j=1}^J \mathbf{e}_{2j} E_{2j}(\tau) + \sum_{j=1}^J \mathbf{b}_{2j+1} B_{2j+1}(\tau)$$





E.-B terms



Remainders

# Short arc representation

LEO orbit can be represented as:

$$\mathbf{r}(\tau) = \bar{\mathbf{r}}(\tau) + \mathbf{d}(\tau) = \bar{\mathbf{r}}(\tau) + \sum_{v=1}^n \mathbf{d}_v \sin(v\pi\tau)$$

Gibbs effect!

or

$$\mathbf{r}(\tau) = \bar{\mathbf{r}}(\tau) + \mathbf{d}(\tau) = \bar{\mathbf{r}}(\tau) + \sum_{j=1}^J \mathbf{e}_{2j} E_{2j}(\tau) + \sum_{j=1}^J \mathbf{b}_{2j+1} B_{2j+1}(\tau)$$

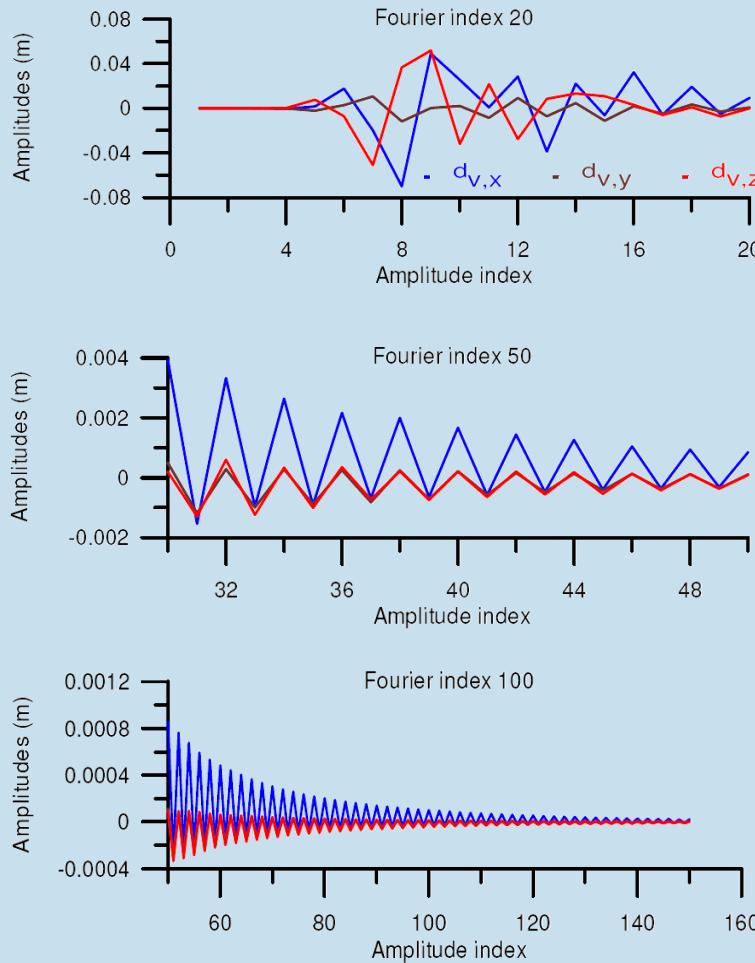
Precision!

Solution?

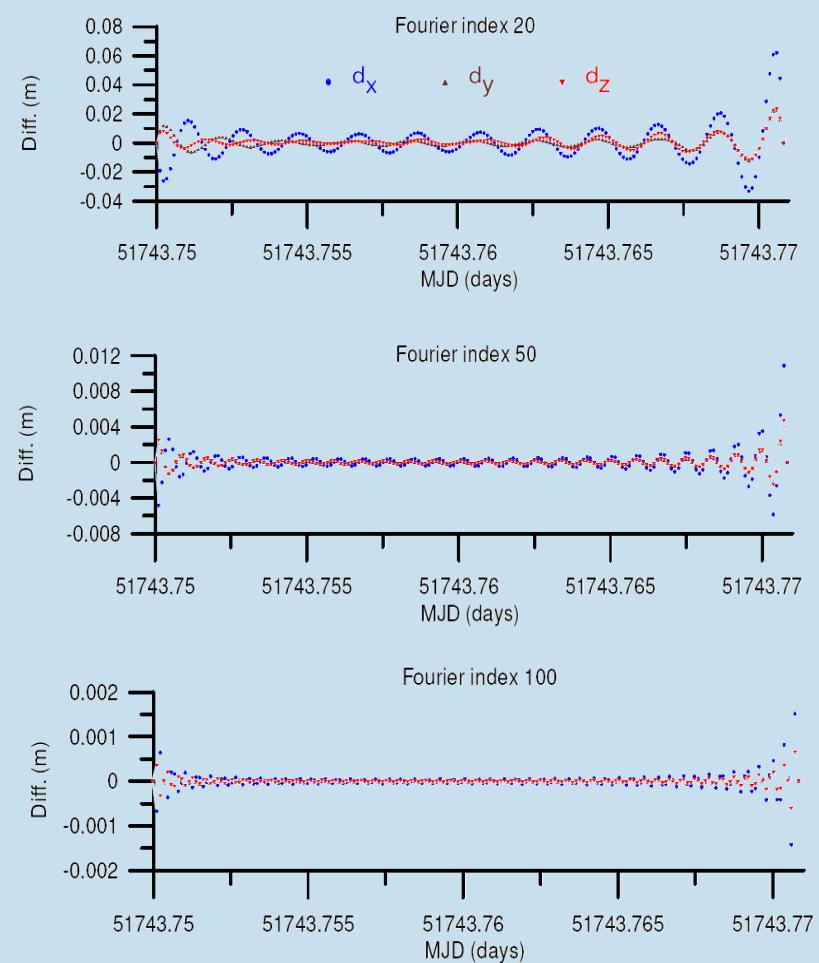
fast convergence!

$$\mathbf{r}(\tau) = \bar{\mathbf{r}}(\tau) + \sum_{j=1}^J \mathbf{e}_{2j} E_{2j}(\tau) + \sum_{j=1}^J \mathbf{b}_{2j+1} B_{2j+1}(\tau) + \sum_{v=1}^{\bar{n}} \bar{\mathbf{d}}_v \sin(v\pi\tau)$$

# Kinematical POD – ellipse mode, J=4



**Amplitudes**



**Remainders**

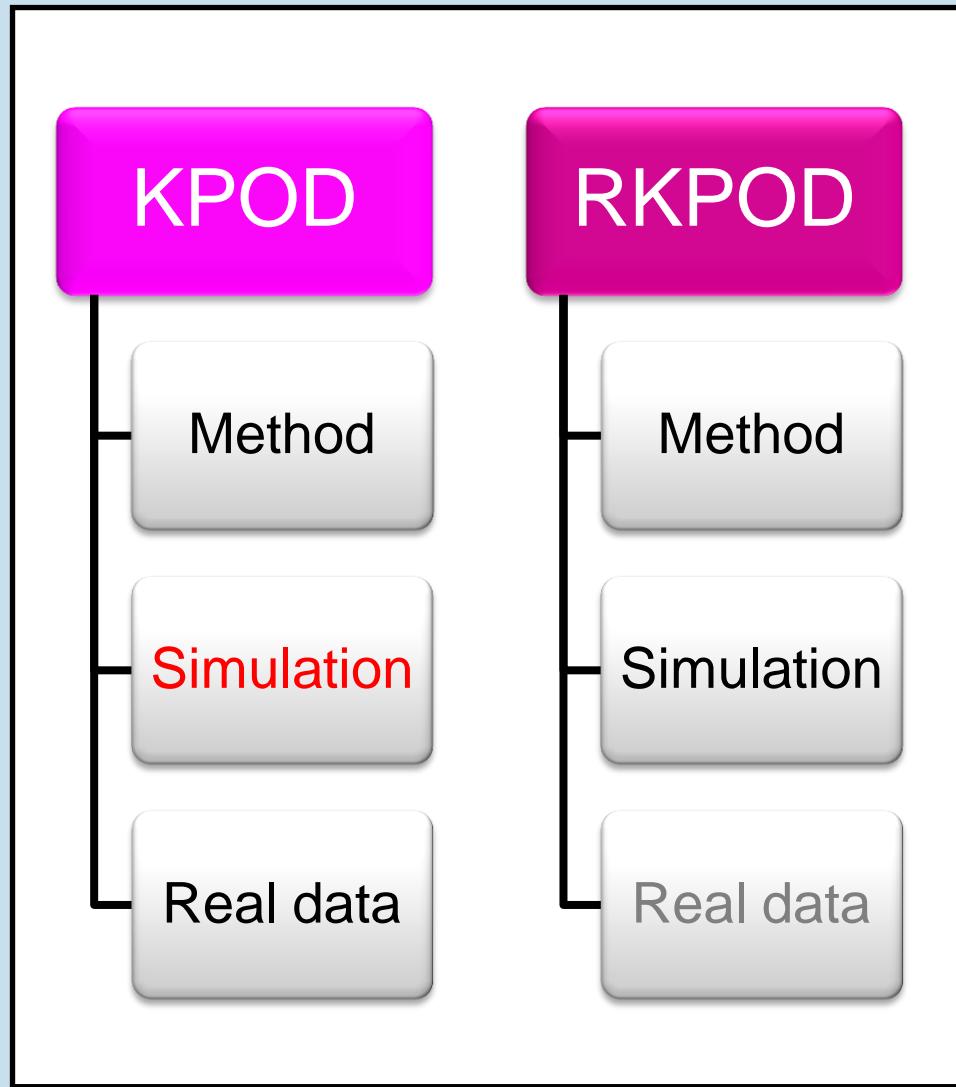
GPOD

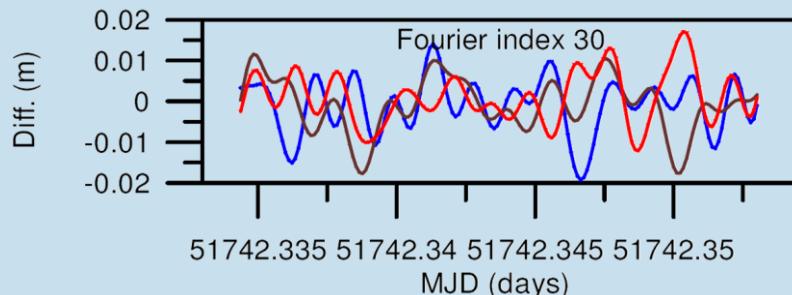
KPOD

RKPOD

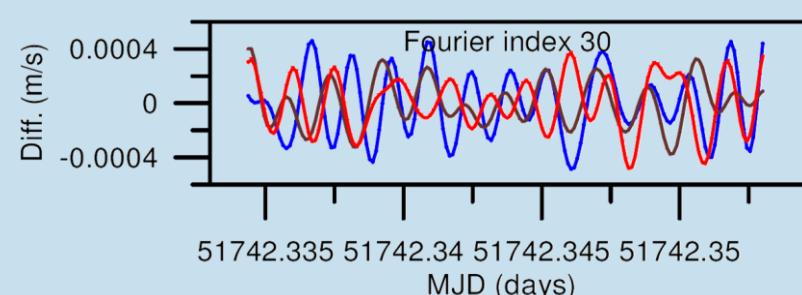
DPOD

17

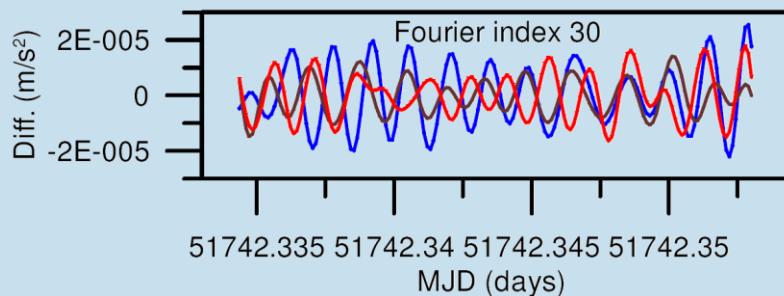




Position differences



Velocity differences



Acceleration differences

RMS

index	Pos.(m)	Vel.(m/s)	Acc.(m/s <sup>2</sup> )
20	0.012644	0.000353	0.000012
30	0.010717	0.000397	0.000018
40	0.011997	0.000463	0.000025
59	0.014737	0.000941	0.000077

Statistical values

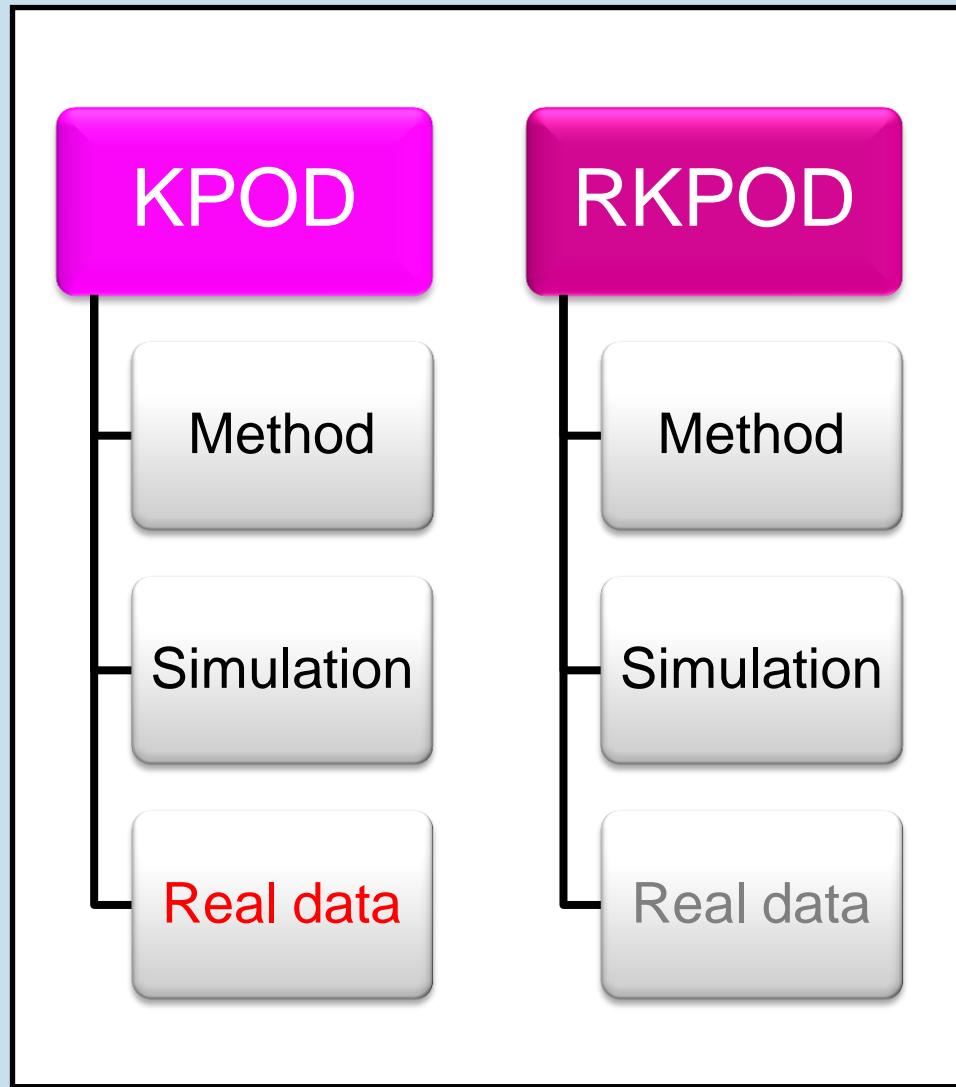
GPOD

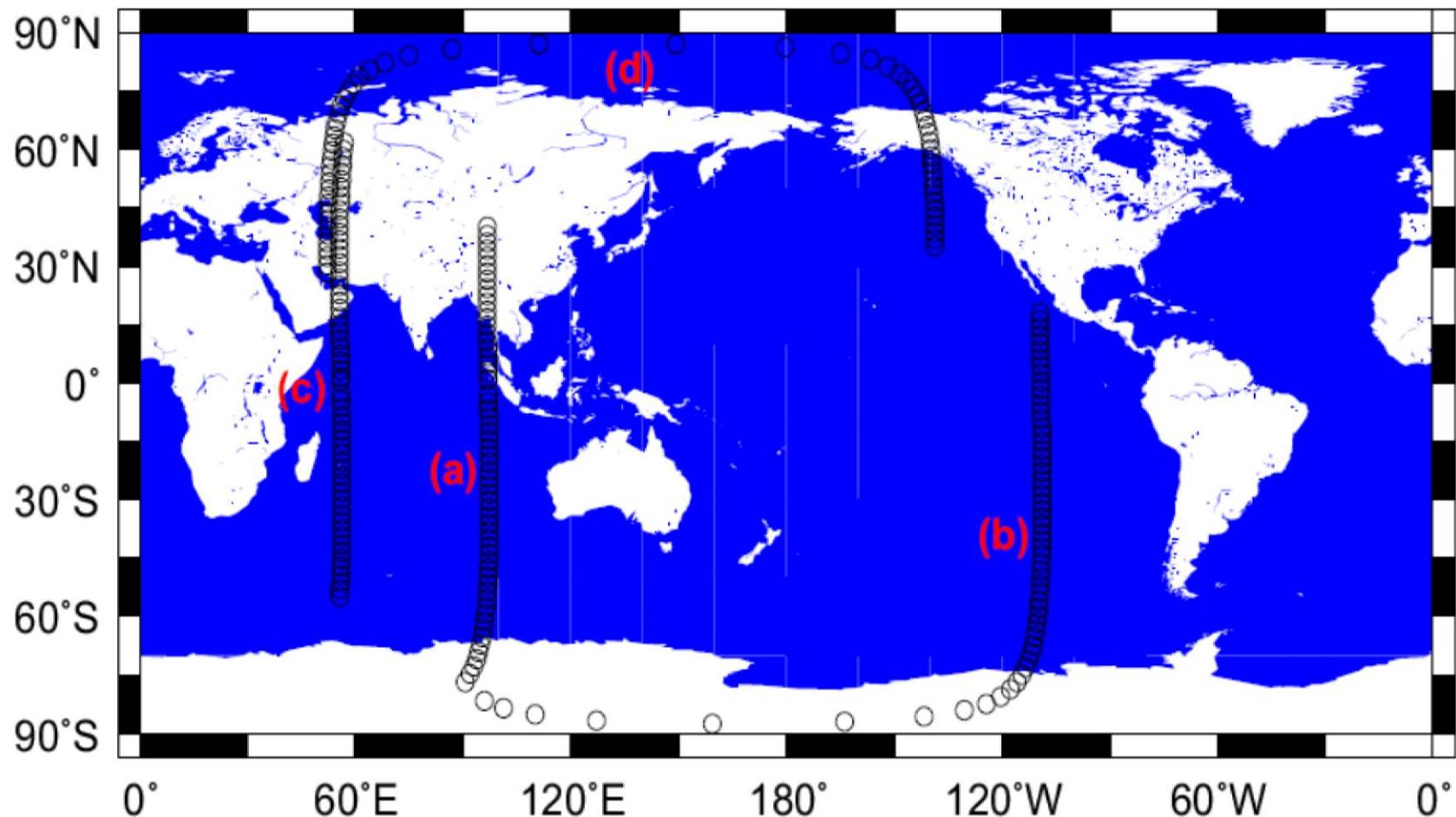
KPOD

RKPOD

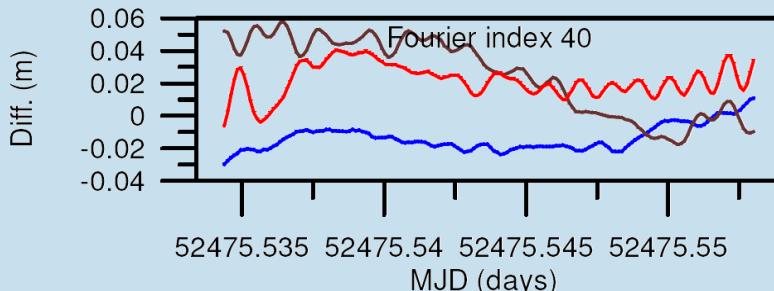
DPOD

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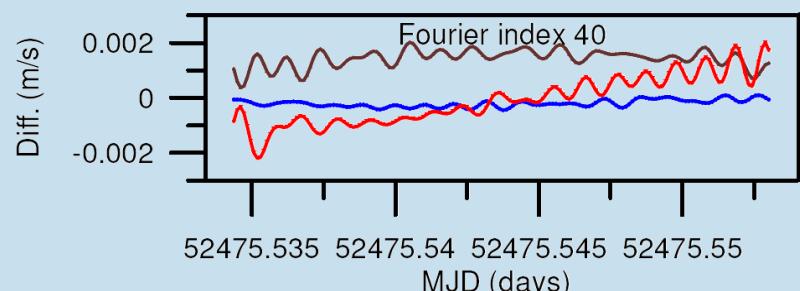




Four short arcs (30 min.) ground track of CHAMP



**IGG - GFZ positions**

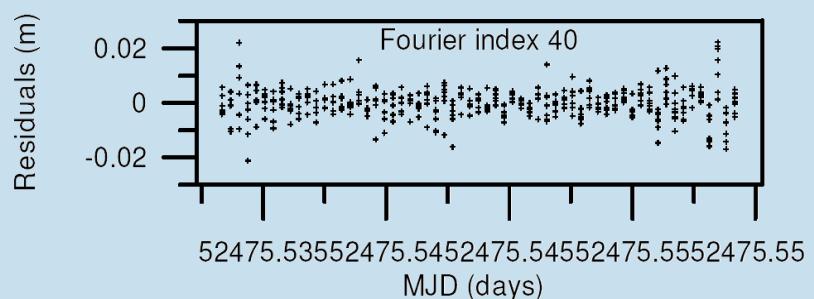


**IGG - GFZ velocities**

RMS

index	Pos.(m)	Vel.(m/s)
20	0.0503	0.0019
30	0.0455	0.0018
40	0.0449	0.0017
59	0.0449	0.0017

**Statistical values**



**GPS-SST residuals**

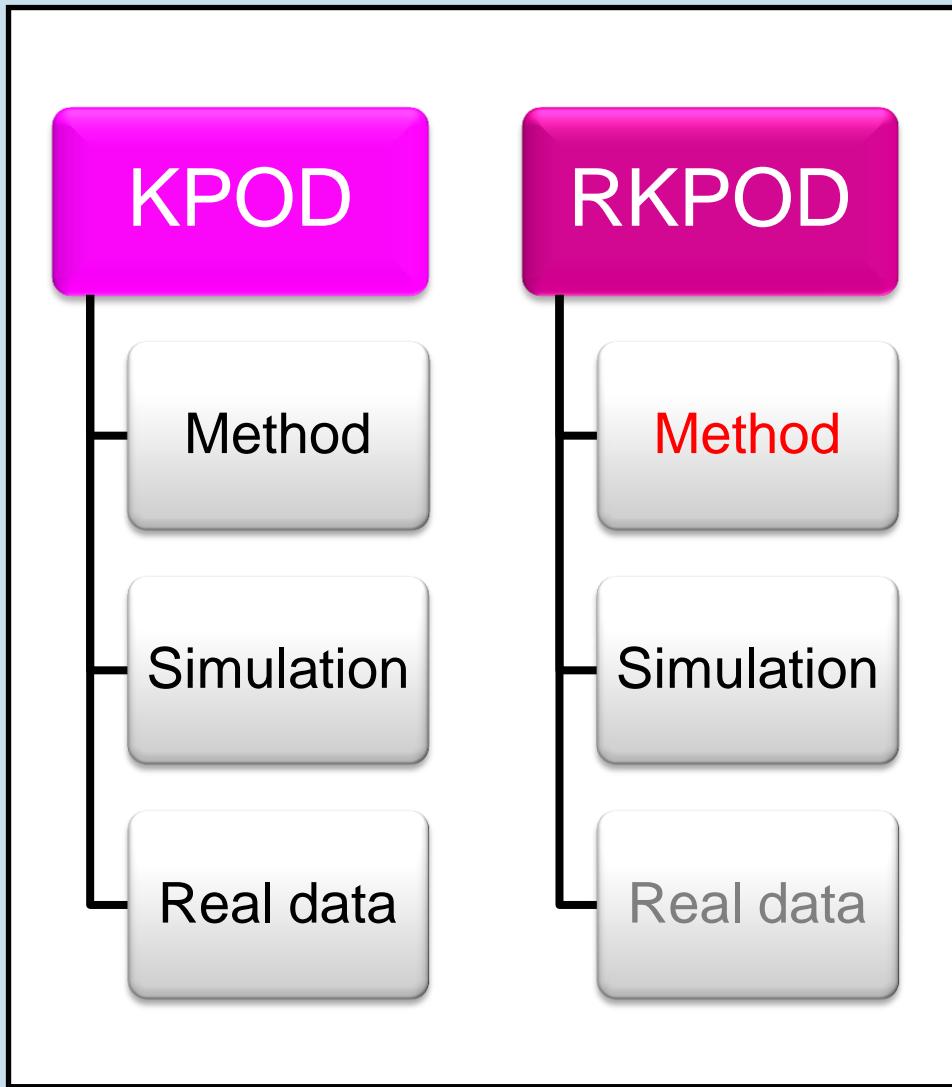
GPOD

KPOD

RKPOD

DPOD

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Dynamical  
info.

$$\mathbf{r} = \bar{\mathbf{r}}(\tau) + \sum_{v=1}^n \tilde{\mathbf{d}}_v \sin(v\pi\tau)$$

$$\tilde{\mathbf{d}}_v = -\frac{2T^2}{v^2\pi^2 - \mu^2} \int_{\tau'=0}^1 \sin(v\pi\tau) \mathbf{a}(\tau'; \mathbf{r}, \dot{\mathbf{r}}) d\tau'$$

- ✓ Introduction of an approximate force function  $(\tilde{\mathbf{d}}_i \cdots \tilde{\mathbf{d}}_j), \mathbf{C}_{(\tilde{\mathbf{d}}_i \cdots \tilde{\mathbf{d}}_j)}$
- ✓ Fixing only some orbit parameters  $(\tilde{\mathbf{d}}_i \cdots \tilde{\mathbf{d}}_j), \mathbf{C}_{(\tilde{\mathbf{d}}_i \cdots \tilde{\mathbf{d}}_j)} \rightarrow 0$
- ✓ Down- or up weighting  $\mathbf{C}_{(\tilde{\mathbf{d}}_1 \cdots \tilde{\mathbf{d}}_n)}$  in relation to  $\mathbf{C}_{(\mathbf{d}_1 \cdots \mathbf{d}_n)}$

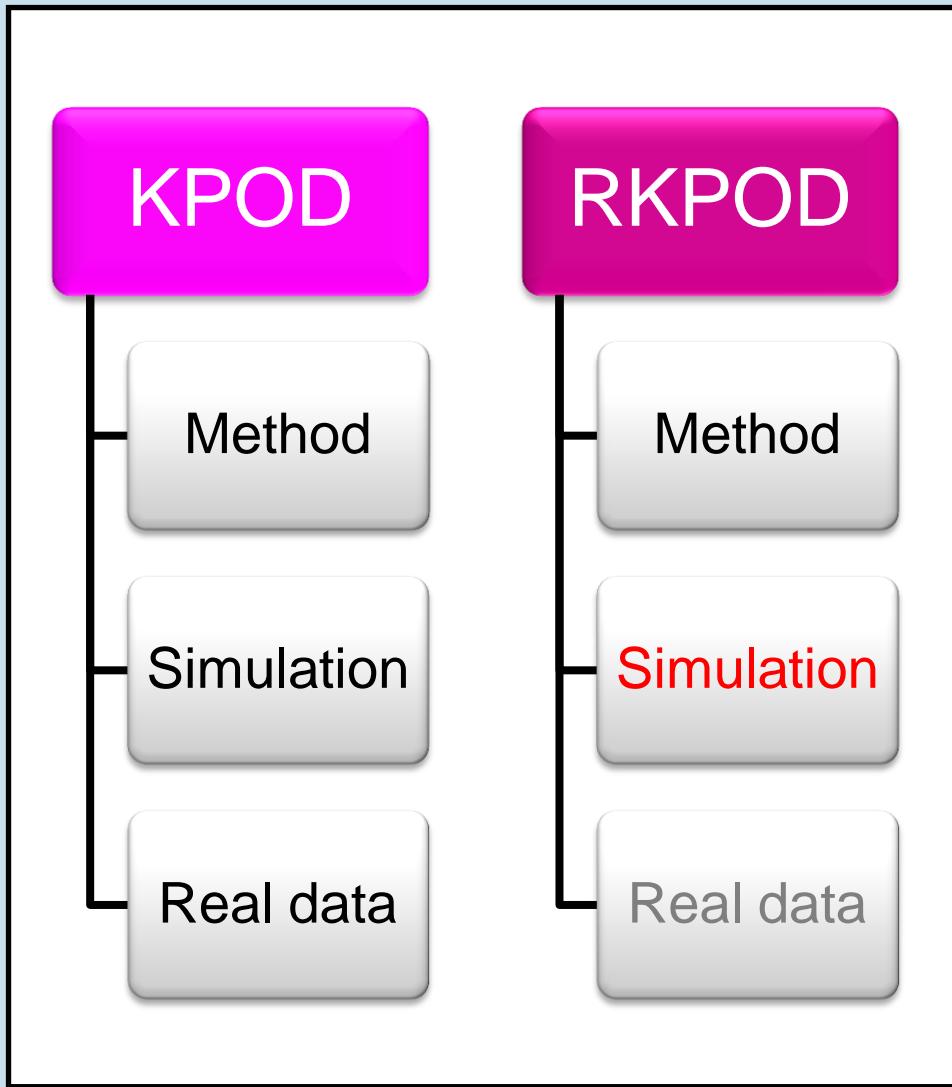
GPOD

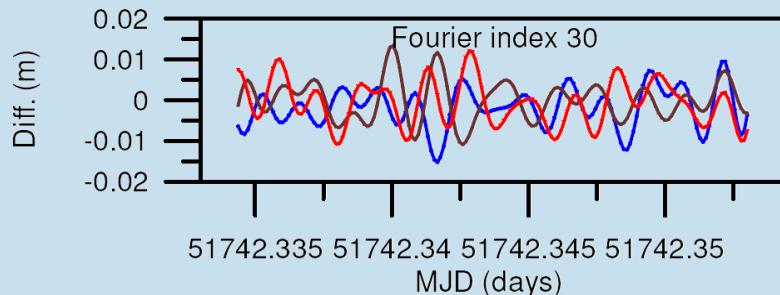
KPOD

RKPOD

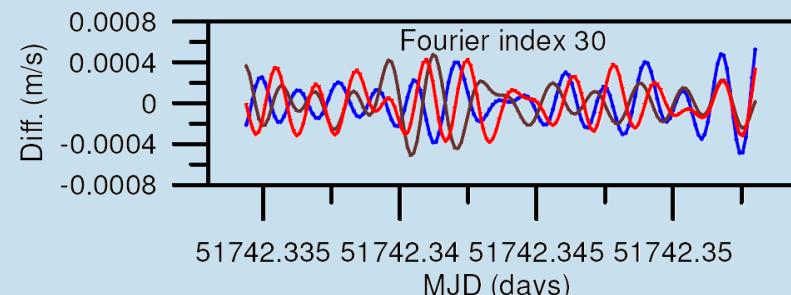
DPOD

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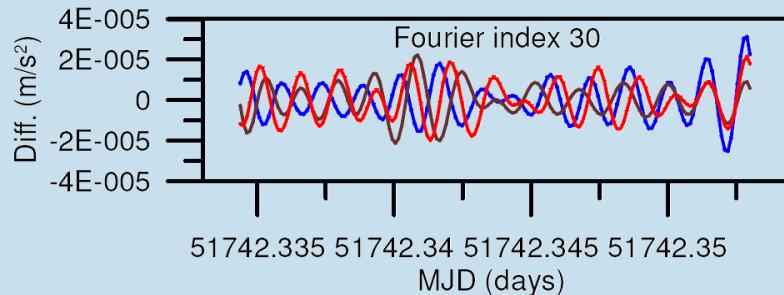




Position differences



Velocity differences



Acceleration differences

RMS

index	Pos.(m)	Vel.(m/s)	Acc.(m/s <sup>2</sup> )
20	0.012831	0.000316	0.000012
30	0.008873	0.000337	0.000016
40	0.014034	0.000402	0.000021
59	0.011553	0.000721	0.000056

Statistical values

- GNSS-LEO satellites configuration and geometrical strength play an important role in POD,
- Kinematical POD can be used to recover the Earth's gravity field model based on the POD methods,
- No gravity field and no force models have been used in the Geometrical and Kinematical modes (**advantage**),
- The proposed kinematical orbit determination method is very flexible. A smooth transition from kinematical to reduced kinematical and finally dynamical or vice-versa is possible.

**Thank you  
for your attention**